



Statistics for Astronomers: Lecture 8, 2020.10.26

Prof. Sundar Srinivasan

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The χ^2 distribution. χ^2 minimisation and interpretation. Reduced χ^2 and caution. The Empirical Rule for normal distributions. The z-score.



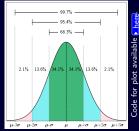
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Probability of obtaining values within 1, 2, or 3 σ of the centre of a normal distribution.

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$P(Z \le 1) \approx 0.68$	$\overline{P(Z>1)=rac{1}{2}\left[1-P(Z \leq 1) ight]}pprox 0.16$
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Therefore also known as the 68–95–99.7 Rule.

 3σ rule of thumb for normal distributions: most (99.7%) of your data is within 3σ of the mean.





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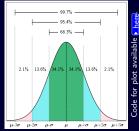
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what is the threshold $z_{\alpha/2}$ such that $P(|Z| > z_{\alpha/2}) < \alpha$? (Two-tailed test) what is the threshold $z_{\alpha/2}$ such that $P(Z > z_{\alpha/2}) < \alpha$? (One-tailed test)

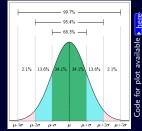
$$P(|Z| > z_{\alpha/2}) = 1 - P(|Z| \le z_{\alpha/2}) = 1 - \operatorname{erf}\left(\frac{z_{\alpha/2}}{\sqrt{2}}\right) \Longrightarrow z_{\alpha/2} = \sqrt{2}\operatorname{erf}^{-1}\left(1 - \alpha\right).$$



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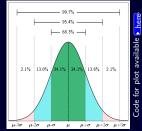
Hypothesis testing. $1 - \alpha$ usually 95%. Observed probability: *p*-value.

Example: pixel with flux 3σ above noise level. p-value: $P(Z > 3) = 0.00135 < \alpha = 0.05$. \implies the detection is statistically significant at the $\alpha = 95\%$ level.



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What if σ is also unknown? We can estimate it from data. Recall: $\widehat{\sigma} = S$ such that

$$S^{2} = \frac{1}{N} \sum_{i=1}^{N} (x_{i} - \mu)^{2} \quad (\mu \text{ known}) \qquad S^{2} = \frac{1}{N-1} \sum_{i=1}^{N} (x_{i} - \overline{x})^{2} \quad (\mu \text{ unknown})$$



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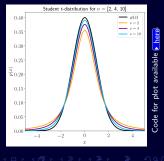
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 with $u=\# ext{dof}=N$ or $N-1.$





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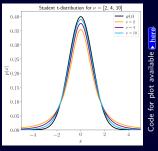
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Symmetric about t = 0, odd moments = 0 (like the Gaussian). Uncertain estimate S for $\sigma \Rightarrow$ more probability in the tails.

$$p_T(t,\nu) \xrightarrow{N \to \infty} \mathscr{N}(0,1) \qquad \qquad \operatorname{Var}[T] = \sqrt{\frac{\nu}{\nu-2}} \xrightarrow{\nu \to \infty} 1.$$





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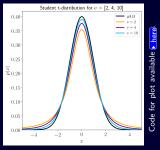
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Use the t-distribution for N < 30.





For small samples (N < 30), we must compute the *t*-equivalent of the *z* statistic in order to determine *t*-scores.

For a Normal random variable X, $T = \frac{X - \mu}{S}$. For any (CLT) sample mean: $T = \frac{\overline{X} - \mu}{S/\sqrt{N}}$



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For $\nu = 4$, using scipy.stats.t.cdf and scipy.stats.norm.cdf, compare the central concentration of T with Z: $P(|T_{\nu=4}| < 1) \approx 0.63; P(|Z| < 1) \approx 0.68$ $P(|T_{\nu=4}| < 2) \approx 0.88; P(|Z| < 2) \approx 0.95$ $P(|T_{\nu=4}| < 3) \approx 0.96; P(|Z| < 3) \approx 0.997$



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Central behaviour quite similar!



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 $\alpha = 0.1$: $t_{\nu=4,\alpha/2} \approx 2.13$; $z_{\alpha/2} \approx 1.64$ using t.ppf(1- $\alpha/2$), norm.ppf(1- $\alpha/2$)
 $\alpha = 0.05$: $t_{\nu=4,\alpha/2} \approx 2.78$; $z_{\alpha/2} \approx 1.96$
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Very different because of behaviour in the tails ("fatter" or "heavier"-tailed distribution)!



More general rule(s) for non-Normal distributions?

If location and scale parameters known:

Definition (Chebyshev's Inequality)

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Two-tailed version, can be modified for asymmetric distributions.

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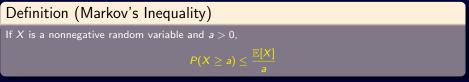
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If μ , σ unknown:



Two-tailed version also exists.





Interval estimates



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References:

"Statistics: A Guide and Reference to the Use of Statistical Methods in the Physical Sciences" - R. J. Barlow.

"Dos and don'ts of reduced chi-squares" - R. Andrae, 2010.

Wall & Jenkins.

